

## Important Notice to Authors

*No further publication processing will occur until we receive your response to this proof.*

Attached is a PDF proof of your forthcoming article in *Physical Review Letters*. The article accession code is LQ17413.

Your paper will be in the following section of the journal: LETTERS — Atomic, Molecular, and Optical Physics

Please note that as part of the production process, APS converts all articles, regardless of their original source, into standardized XML that in turn is used to create the PDF and online versions of the article as well as to populate third-party systems such as Portico, Crossref, and Web of Science. We share our authors' high expectations for the fidelity of the conversion into XML and for the accuracy and appearance of the final, formatted PDF. This process works exceptionally well for the vast majority of articles; however, please check carefully all key elements of your PDF proof, particularly any equations or tables.

Figures submitted electronically as separate files containing color appear in color in the online journal.

However, all figures will appear as grayscale images in the print journal unless the color figure charges have been paid in advance, in accordance with our policy for color in print (<https://journals.aps.org/authors/color-figures-print>).

### Specific Questions and Comments to Address for This Paper


The numbered items below correspond to numbers in the margin of the proof pages pinpointing the source of the question and/or comment. The numbers will be removed from the margins prior to publication.

- 1 Except for the term "and/or," the use of the slash is discouraged between words and abbreviations, as the intent of the solidus is ambiguous. Several possibilities for its meaning exist, among them "and," "or," "and/or," and "plus." We ask that more precise, and therefore more meaningful, conjunctions be used. For terms that are diagrammatically opposed, we use a hyphen (e.g., liquid-solid interface, vacancy-acceptor interface).
- 2 It is journal style to break down fractions in superscripts and subscripts. Please check that fractions and added bracketing in Eqs. (6),(7),(9) appear acceptably.
- 3 Please write out FWM. Was FWHM intended?
- 4 Please review the funding information section of the proof's cover letter and respond as appropriate. We must receive confirmation that the funding agencies have been properly identified before the article can publish.
- 5 NOTE: External links, which appear as blue text in the reference section, are created for any reference where a Digital Object Identifier (DOI) can be found. Please confirm that the links created in this PDF proof, which can be checked by clicking on the blue text, direct the reader to the correct references online. If there is an error, correct the information in the reference or supply the correct DOI for the reference. If no correction can be made or the correct DOI cannot be supplied, the link will be removed.
- 6 A check of online databases revealed a possible error in Ref. [30]. The year has been changed from '2016' to '2017'. Please confirm this is correct.

### Titles in References

The editors now encourage insertion of article titles in references to journal articles and e-prints. This format is optional, but if chosen, authors should provide titles for *all* eligible references. If article titles remain missing from eligible references, the production team will remove the existing titles at final proof stage.

### ORCIDs

Please follow any ORCID links () after the authors' names and verify that they point to the appropriate record for each author.

## Funding Information

Information about an article's funding sources is now submitted to Crossref to help you comply with current or future funding agency mandates. Crossref's Open Funder Registry (<https://www.crossref.org/services/funder-registry/>) is the definitive registry of funding agencies. Please ensure that your acknowledgments include all sources of funding for your article following any requirements of your funding sources. Where possible, please include grant and award ids. Please carefully check the following funder information we have already extracted from your article and ensure its accuracy and completeness:

- Ministry of Education - Singapore, FundRef ID <http://dx.doi.org/10.13039/501100001459> (Republic of Singapore/SG)

## Other Items to Check

- Please note that the original manuscript has been converted to XML prior to the creation of the PDF proof, as described above. Please carefully check all key elements of the paper, particularly the equations and tabular data.
- Title: Please check; be mindful that the title may have been changed during the peer-review process.
- Author list: Please make sure all authors are presented, in the appropriate order, and that all names are spelled correctly.
- Please make sure you have inserted a byline footnote containing the email address for the corresponding author, if desired. Please note that this is not inserted automatically by this journal.
- Affiliations: Please check to be sure the institution names are spelled correctly and attributed to the appropriate author(s).
- Receipt date: Please confirm accuracy.
- Acknowledgments: Please be sure to appropriately acknowledge all funding sources.
- References: Please check to ensure that titles are given as appropriate.
- Hyphenation: Please note hyphens may have been inserted in word pairs that function as adjectives when they occur before a noun, as in “x-ray diffraction,” “4-mm-long gas cell,” and “*R*-matrix theory.” However, hyphens are deleted from word pairs when they are not used as adjectives before nouns, as in “emission by x rays,” “was 4 mm in length,” and “the *R* matrix is tested.”  
Note also that Physical Review follows U.S. English guidelines in that hyphens are not used after prefixes or before suffixes: superresolution, quasiequilibrium, nanoprecipitates, resonancelike, clockwise.
- Please check that your figures are accurate and sized properly. Make sure all labeling is sufficiently legible. Figure quality in this proof is representative of the quality to be used in the online journal. To achieve manageable file size for online delivery, some compression and downsampling of figures may have occurred. Fine details may have become somewhat fuzzy, especially in color figures. The print journal uses files of higher resolution and therefore details may be sharper in print. Figures to be published in color online will appear in color on these proofs if viewed on a color monitor or printed on a color printer.
- Overall, please proofread the entire *formatted* article very carefully. The redlined PDF should be used as a guide to see changes that were made during copyediting. However, note that some changes to math and/or layout may not be indicated.

## Ways to Respond

- **Web:** If you accessed this proof online, follow the instructions on the web page to submit corrections.
- **Email:** Send corrections to [aps-robot@luminad.com](mailto:aps-robot@luminad.com). Include the accession code LQ17413 in the subject line.
- **Fax:** Return this proof with corrections to +1.855.808.3897.

**If You Need to Call Us**

You may leave a voicemail message at +1.855.808.3897. Please reference the accession code and the first author of your article in your voicemail message. We will respond to you via email.

## Spectral Compression of Narrowband Single Photons with a Resonant Cavity

Mathias A. Seidler,<sup>1</sup> Xi Jie Yeo,<sup>2</sup> Alessandro Cerè<sup>1</sup>,<sup>1</sup> and Christian Kurtsiefer<sup>1,2,\*</sup>

<sup>1</sup>Centre for Quantum Technologies, National University of Singapore, 3 Science Drive 2, Singapore 117543

<sup>2</sup>Department of Physics, National University of Singapore, 2 Science Drive 3, Singapore 117551

(Received 19 March 2020; accepted 7 October 2020)

We experimentally demonstrate a spectral compression scheme for heralded single photons with narrow spectral bandwidth around 795 nm, generated through four-wave mixing in a cloud of cold <sup>87</sup>Rb atoms. The scheme is based on an asymmetric cavity as a dispersion medium and a simple binary phase modulator, and can be, in principle, without any optical losses. We observe a compression from 20.6 MHz to less than 8 MHz, almost matching the corresponding atomic transition.

DOI:

*Introduction.*—Efficient atom-light interactions at the single quantum level is at the core of several proposals for storing, processing, and relaying quantum information [1–4]. Many of these schemes require single “flying” photons to match the spectrum of atomic transitions [5–8]. Single photons can be emitted from trapped ions [9,10], atoms [11–13], or solid-state systems [14–16]. However, the spectral width of the generated photons may not always match the spectral width of the receiving systems. Therefore, methods to engineer the photon spectrum may be required.

The simplest method for this is to passively filter the spectrum of bright broadband sources [17,18], with a sometimes significant reduction of brightness, making photon-atom interaction experiments that require a high interaction rate [19,20] difficult. More advanced methods to manipulate the spectrum of single photon sources to match that of atomic transitions include restricting the spectral mode of emitters with cavities [11,16,21,22], or using electromagnetically induced transparency in atomic ensembles [23,24] in the source mechanism altogether. As spectral filtering or engineering of the photon generation mechanism may not always be possible, it would be desirable to modify the spectrum of a given photon source while maintaining the brightness. To our knowledge, the only experiments to modify the photon spectrum of narrowband single photons use gradient echo quantum memories [25,26]. However, this was only demonstrated for photons with spectral bandwidths narrower than atomic absorption linewidths.

Here, we demonstrate an alternative technique that compresses the spectral bandwidth of single photons with a spectral bandwidth a few times broader than the corresponding atomic absorption linewidth, while in principle, maintaining the photon rates. The technique is based on the ideas of time lenses invented for temporal imaging [27,28], where the temporal and spectral characteristics of ultrafast electromagnetic pulses [29–31] are manipulated. It turns

out that single photon states can be manipulated in a similar way, complementing the techniques for lossless temporal envelope manipulation of narrowband single photon states demonstrated in [32,33].

Spectral compression of single photon wave packets is achieved in two steps. First, the wave packet is spread out in time such that the width of its envelope is compatible with a narrow spectrum; this can be done using a dispersive element that spreads out different frequency components of the wave packet in time, effectively generating a chirped wave packet. In the second step, a time-dependent phase shift is applied. This step changes the spectral energy distribution of the wave packet.

Previous time-lens-based spectral compression schemes were performed on ultrashort pulses, using optical fibers or diffraction gratings as dispersive elements [29–31]. The suitability of a dispersive element for a spectral compression scheme is related to the spectral bandwidth of the optical pulse. The bandwidths of the ultrashort pulses used in previous time-lens-based spectral compression schemes are typically on the order of 0.1, ..., 1 THz, and the length of optical fibers used to generate a significant temporal broadening of these pulses are on the order of 0.1, ..., 10 km. However, photonic wave packets interacting with single emitters like atoms or molecules have spectral bandwidths on the order of MHz, which would require optical fibers on the order of 10<sup>8</sup> m to generate a suitable temporal broadening for spectral compression. Transmitting light through such a long fiber would not only be impractical, but also prohibitively lossy. Similarly, currently available gratings would not be able to significantly disperse photonic wave packets with bandwidths of a few MHz. We overcome this problem by using the dispersive properties of an optical cavity instead. While the dispersion in optical cavities can be much larger, the process then requires a different time-dependent phase shift in the second step to complete the spectral compression process.

89 *Theory.*—To understand the spectral compression  
 90 scheme, we start with an initial single photon wave packet,  
 91 described by an envelope  $|\psi(t)|^2$  of its intensity in time, and  
 92 its corresponding power spectrum  $|\Psi(\omega; \omega_0, \Gamma_p)|^2$ , con-  
 93 nected by the Fourier transform  $F$ :  $\Psi(\omega) = F[\psi(t)]$ . The  
 94 nearly monochromatic wave packet shall be characterized  
 95 by a central frequency  $\omega_0$  and a spectral width  $\Gamma_p$ . The  
 96 spreading out of the wave packet in time is accomplished  
 97 **1** by reflection off an asymmetric cavity, with an input-output  
 98 coupler with a low transmission, and a second high-  
 99 reflective mirror, similar to the setup used in [34].

100 If the losses in the cavity are negligible compared to the  
 101 transmission of the coupling mirror, and the cavity line-  
 102 width  $\Gamma_c$  and photon bandwidth  $\Gamma_p$  are much smaller than  
 103 free spectral range of the cavity, the action of the cavity to a  
 104 wave packet near its resonance  $\omega_c$  can be described by a  
 105 transfer function

$$C(\omega; \omega_c, \Gamma_c) \approx -\frac{\Gamma_c + i2(\omega - \omega_c)}{\Gamma_c - i2(\omega - \omega_c)}, \quad (1)$$

106 which modifies the incoming spectral wave packet  
 107  $\Psi(\omega; \omega_0, \Gamma_p)$  to a new one,

$$\Psi'(\omega; \Delta\omega, \Gamma_c, \Gamma_p) = \Psi(\omega; \omega_0, \Gamma_p)C(\omega; \omega_c, \Gamma_c), \quad (2)$$

108 where  $\Delta\omega = \omega_0 - \omega_c$  is the detuning between the wave  
 109 packet and the cavity resonance. For a lossless cavity, this  
 110 wave packet has the same power spectrum as  $\Psi(\omega)$  because  
 111  $|C(\omega; \omega_c, \Gamma_c)|^2 = 1$ . The temporal envelope of the reflected  
 112 wave packet, obtained through the inverse Fourier trans-  
 113 form  $F^{-1}$ ,  
 114  
 115

$$\psi'(t; \Delta\omega, \Gamma_c, \Gamma_p) = F^{-1}[\Psi'(\omega; \Delta\omega, \Gamma_c, \Gamma_p)], \quad (3)$$

116 is now broader in time, and has acquired a time-dependent  
 117 phase  $\phi'(t; \Delta\omega, \Gamma_p, \Gamma_c)$ .

118 Similar to Fourier-transform limited pulses, where the  
 119 time-bandwidth product is minimized by a frequency-  
 120 independent spectral phase, we can reduce the spectral  
 121 bandwidth of the heralded single photon by removing any  
 122 time-dependent phase. This is done by applying a time-  
 123 dependent phase shift  
 124

$$\phi_e(t) = -\phi'(t; \Delta\omega, \Gamma_p, \Gamma_c), \quad (4)$$

125 resulting in the spectrally compressed wave packet

$$\psi''(t; \Delta\omega, \Gamma_p, \Gamma_c) = \psi'(t; \Delta\omega, \Gamma_p, \Gamma_c)e^{i\phi_e(t)}. \quad (5)$$

126 To quantify the compression, we compare the spectral  
 127 widths before and after the compression obtained from the  
 128 respective power spectrum  $|\psi''(\omega; \Delta\omega, \Gamma_p, \Gamma_c)|^2$  obtained  
 129 through a Fourier transform of Eq. (5).

130 We now consider the specific case of a heralded single  
 131 photon emerging from an atomic cascade decay, where we  
 132 intend to compress the idler photon (see inset of Fig. 2).  
 133 Detection of a signal photon projects the field in the idler  
 134 mode into the heralded state **2**

$$\psi(t) = \sqrt{\Gamma_p}e^{-\Gamma_p/[2(t-t_0)]}\Theta(t-t_0), \quad (6)$$

135 where  $t_0$  and  $t$  are the detection times of the signal and idler  
 136 photons, respectively. The exponential decay with the  
 137 constant  $\Gamma_p$  is a characteristic of the spontaneous process,  
 138 while the Heaviside step function  $\Theta$  is a consequence of the  
 139 well-defined time order of the cascade decay process. For  
 140 simplicity, we set  $t_0 = 0$ .

141 This temporal profile corresponds to a Lorentzian power  
 142 spectrum for the idler photons, and its bandwidth is  
 143 described by the full width at half maximum  $\Gamma_p$ , which  
 144 also corresponds to the spectral window containing 50% of  
 145 the total pulse energy. However, the compressed spectrum  
 146  $|F^{-1}[\psi''(t)]|^2$  has multiple maxima, and is distinctly differ-  
 147 ent from distributions where the full width at half maximum  
 148 naturally quantifies the bandwidth. Hence, we instead define  
 149 bandwidth as the smallest spectral width containing 50% of  
 150 the total pulse energy, as this definition of bandwidth is  
 151 compatible for both a Lorentzian and a generic spectrum.

152 To obtain the optimal cavity parameters, we numerically  
 153 minimize the bandwidth of the compressed photon spec-  
 154 trum. We find that the maximal compression is achieved by  
 155 a resonant cavity  $\Delta\omega = 0$  with a bandwidth of  $\Gamma_c \approx \Gamma_p/4$ .  
 156 Under these conditions, the compressed single photon time  
 157 envelope can be written as  
 158  
 159  
 160  
 161

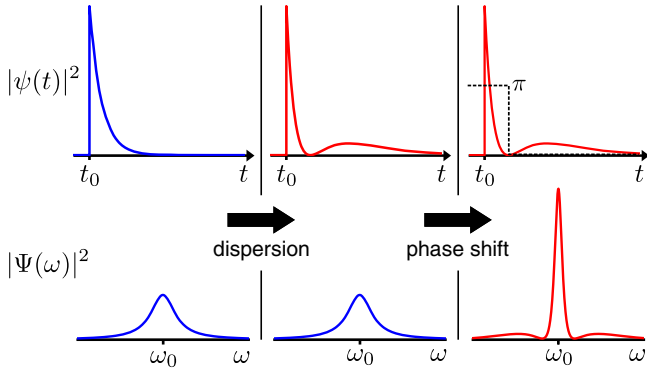
$$\psi''(t) = e^{-i\phi'(t)} \sqrt{\Gamma_p} \frac{2\Gamma_c e^{-(\Gamma_c/2)t} - (\Gamma_p + \Gamma_c) e^{-(\Gamma_p/2)t}}{\Gamma_p - \Gamma_c} \Theta(t), \quad (7)$$

162 with a phase function

$$\phi'(t) = \pi\Theta\left(t - 2\frac{\log\left(\frac{\Gamma_p + \Gamma_c}{2\Gamma_c}\right)}{\Gamma_p - \Gamma_c}\right). \quad (8)$$

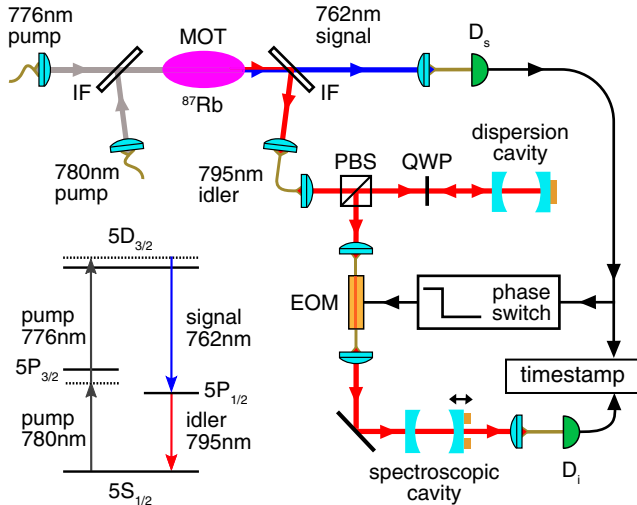
163 This is a step function changing the phase by  $\pi$ , with the  
 164 transition occurring at the minimum of the dispersed  
 165 photon's temporal intensity profile. The narrowest band-  
 166 width achievable with compression based on an asymmet-  
 167 ric cavity with this strategy is  $\sim 0.3\Gamma_p$ ; the temporal  
 168 envelopes and power spectra shown in Fig. 1 correspond  
 169 to this choice.

170 *Experiment.*—Details of the actual experiment are shown  
 171 in Fig. 2. We generate the time-ordered photon pairs by  
 172 four-wave mixing in a cold ensemble of  $^{87}\text{Rb}$  atoms in a  
 173 cascade level scheme [12]. Pump beams at 780 and 776 nm  
 174 excite atoms from the  $5S_{1/2}, F = 2$  ground level to the  
 175  $5D_{3/2}, F = 3$  level via a two-photon transition. The 762 nm  
 176 (signal) and 795 nm (idler) photon pairs emerge from a  
 177  
 178



F1:1 FIG. 1. Concept of spectral compression. The top row shows  
 F1:2 temporal intensity profiles  $|\psi(t)|^2$  in various stages of the spectral  
 F1:3 compression, the bottom row the corresponding power spectra  
 F1:4  $|\Psi(\omega)|^2$ . The initial pulse is dispersed by a cavity, leading to a  
 F1:5 new temporal shape, but an unchanged spectrum. An electro-  
 F1:6 optical modulator (EOM) manipulates the phase  $\phi'(t)$  of the pulse  
 F1:7 which leads to a narrower spectrum.

179 cascade decay back to the ground level, and are coupled to  
 180 single mode fibers. Phase matching is ensured with all four  
 181 modes propagating collinearly in the same direction. The  
 182 two pumps have a focus in the cloud with a beam waist of  
 183 about  $400 \mu\text{m}$ . The 780 nm pump is 55 MHz blue detuned  
 184 from the  $5S_{1/2}, F = 2$  to  $5P_{3/2}, F = 3$  transition and has an  
 185 optical power of 0.25 mW. The 776 nm pump has an optical  
 186 power of 11.4 mW, and is tuned such that the two-photon  
 187 transition to the  $5D_{3/2}, F = 3$  state is 5 MHz blue detuned.  
 188 When the excited atoms decay via the  $5D_{1/2}, F = 2$  state  
 189 back into the initial ground state, photons with a wave-  
 190 length of 762 nm and 795 nm photon are emitted [12].



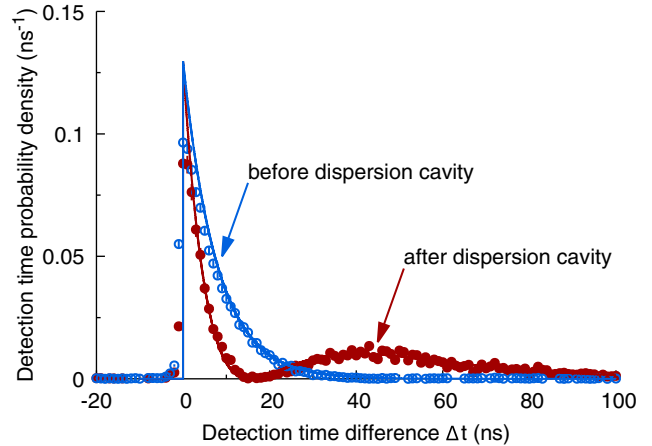
F2:1 FIG. 2. Schematic setup for generation and spectral compression  
 F2:2 of heralded single photons.  $D_{S,I}$ : single-photon detectors;  
 F2:3 EOM: electro-optical modulator; PBS: polarizing beam splitter;  
 F2:4 QWP: quarter-wave plate; IF: interference filter. Inset: energy  
 F2:5 level scheme for four-wave mixing in  $^{87}\text{Rb}$ .

After suppressing residual pump light and separating  
 signal and idler photons into different modes, we collect  
 them into single mode fibers. The 762 nm signal photons  
 are detected with an avalanche photo diode and herald the  
 presence of 795 nm idler photons. The time correlation  
 between the detection in the signal and idler modes (open  
 circles in Fig. 3) correspond to the envelope  $|\psi(t)|^2$  of the  
 intensity in time.

We measure the initial power spectrum of the wave  
 packet (open circles in Fig. 4) by correlating it with a the  
 photon rate transmitted through a Fabry-Perot cavity (FP)  
 with linewidth  $\Gamma_{\text{FP}} \approx 2\pi \times 2.6 \text{ MHz}$ . The transmission is  
 recorded at different detunings of the cavity from the  
 atomic resonance. The observed spectrum was observed  
 to have a full width at half maximum of 20(2) MHz, wider  
 than the atomic linewidth of 6 MHz due to collective  
 emission effects in the cloud [35,36].

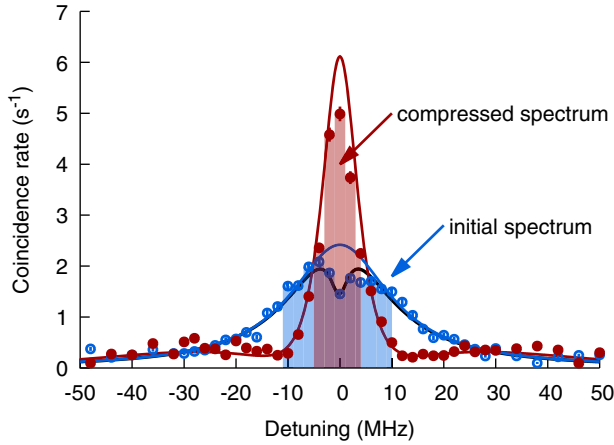
The 795 nm idler photons are then coupled to the  
 dispersion cavity, with a coupling mirror of nominal  
 reflectivity  $R_1 = 0.97$ , and a high reflector with  $R_2 =$   
 0.9995 separated by 10.1 cm, corresponding to a free  
 spectral range of 1.48 GHz, and a measured linewidth  
 $\Gamma_c \approx 2\pi \times 7.3 \text{ MHz}$ . A Pound-Drever-Hall frequency lock  
 keeps the cavity resonant to the central frequency of the  
 photons throughout the experiment. The measured time  
 envelope of the single photon wave packet after dispersion  
 is shown as filled dots in Fig. 3.

The spectral compression is completed by applying the  
 temporal phase of Eq. (8), in the form of a phase switch  
 synchronized to the photon passage through a fiber con-  
 nected electro-optical modulator (EOM). Since the idler  
 photon is heralded, we use the detection of the signal  
 photon as a reference signal for triggering the phase switch



F3:1 FIG. 3. Detection time distribution for the heralded photon  
 F3:2 before (open blue circles) and after (filled red dots) the dispersion  
 F3:3 cavity. We fit an exponential decay Eq. (6) to the initial time  
 F3:4 correlation (blue solid line), from which we infer the photon  
 F3:5 bandwidth  $\Gamma_p$ . The simulated temporal profile after the photon  
 F3:6 passed through the dispersion cavity [red line, calculated from  
 F3:7 Eq. (2)], matches our experimental data (filled dots) well.





F4:1 FIG. 4. Spectral profile of heralded photons before (blue) and  
 F4:2 after (red) spectral compression, obtained by measuring the  
 F4:3 photon transmission rate through the Fabry-Perot cavity at  
 F4:4 different cavity detunings. The solid lines are calculated from  
 F4:5 Eq. (5), with  $\Gamma_p$  inferred from the temporal envelope measure-  
 F4:6 ment of the photons from the source, and  $\Gamma_c$  by experimentally  
 F4:7 characterizing the cavity bandwidth. Shaded areas cover 50% of  
 F4:8 the total power for each spectrum.

224 after an appropriate time delay. Our dispersed photon is  
 225 approximately 80 ns long, which corresponds to a spatial  
 226 spread of 16 m in a fiber with a refractive index of 1.5. The  
 227 phase modulator has an active length of 90 mm, so at any  
 228 instant only a small part of the photon resides inside the  
 229 EOM, and we are able to modulate the two parts of the  
 230 photon with different phases. The correct timing  
 231 of the phase modulation is ensured by measuring the  
 232 length of the fibers and the electric signal lines with a  
 233 timing uncertainty  $< 0.5$  ns, on par with the electrical rise  
 234 time of the phase change signal. This order of timing  
 235 uncertainty can be tolerated, since the majority of second  
 236 photon wave packet part (Fig. 3, blue line from 15 to  
 237 100 ns) gets a phase shift of  $\pi$ . The phase flip is applied  
 238 right after the first part of the dispersed photon exits the  
 239 modulator, and the second part starts to propagate through  
 240 it. This timing is indicated as the dashed line in Fig. 1. We  
 241 finally measure the compressed photon spectrum by again  
 242 recording the photon transmission rate through the Fabry-  
 243 Perot cavity, shown as filled dots in Fig. 4.

244 To obtain an initial photon bandwidth  $\Gamma_p$ , we fit the  
 245 decaying exponential term in Eq. (6) to the observed  
 246 coincidence probability (open circles in Fig. 3). The solid  
 247 red line in Fig. 3 corresponds to an expected temporal  
 248 profile of the photon after the dispersion cavity, calculated  
 249 from Eq. (7), with an inferred photon bandwidth  
 250  $\Gamma_p = 2\pi \times 20.6(2)$  MHz obtained from the fit of the initial  
 251 photon shape, and the cavity linewidth  $\Gamma_c \approx 2\pi \times 7.3$  MHz  
 252 measured earlier. The observed temporal envelope after the  
 253 dispersion cavity (full dots in Fig. 3) agrees very well with  
 254 the expected profile.

255 The measured spectral profiles before and after com-  
 256 pression are shown in Fig. 4. The spectrum of the

257 uncompressed photons (open circles) exhibits a dip around  
 258 the central frequency, which was also observed without the  
 259 compression optics. We attribute this to reabsorption of the  
 260 generated photons by the atomic cloud. We model this  
 261 spectrum  $S(\omega)$  by considering two processes: First, we  
 262 consider the spectrum  $P(\omega)$  of the photon emitted by the  
 263 atomic cloud which can be obtained by the product of the  
 264 spectrum of the photon produced by our FWM process **3**  
 $L(\omega; \Gamma_p) = (2/\pi)(\Gamma_p/4\omega^2 + \Gamma_p^2)$  and an absorption term  
 265 describing the attenuation of the photon by our atomic  
 266 cloud of optical density OD  
 267

$$P(\omega; \text{OD}, \Gamma_p, \Gamma_a) = AL(\omega; \Gamma_p)e^{-\text{OD}[\Gamma_a^2/(4\omega^2 + \Gamma_a^2)]}. \quad (9)$$

269 The scaling factor  $A$  is used to account for the detected  
 270 coincidence rate, and  $\Gamma_a$  is the spectral width of  
 271 the absorption feature. From our fit, we extract  
 272  $\Gamma_a = 2\pi \times 1.38(1.6)$  MHz, which does not correspond to  
 273 the absorption linewidth of the corresponding atomic  
 274 transition ( $2\pi \times 5.7$  MHz) of the  $5S_{1/2} \rightarrow 5P_{1/2}$  transition.  
 275 Further work is necessary to understand this observation.  
 276 The inferred photon bandwidth  $\Gamma_p = 2\pi \times 20.6(2)$  MHz  
 277 was determined from the photon coincidence time correla-  
 278 tion (Fig. 3). Second, we consider the effect of the  
 279 Fabry-Perot cavity used to sample the spectral profile by  
 280 convolution the above result with  $L'(\omega; \Gamma_{\text{FP}}) =$   
 $[\Gamma_{\text{FP}}^2/(4\omega^2 + \Gamma_{\text{FP}}^2)]$  to model the observed spectrum:  
 281

$$S(\omega) = (P * L')(\omega). \quad (10)$$

282 We fit the model  $S(\omega)$  (Fig. 4, black line) to the exper-  
 283 imental data corresponding to the spectral profile of the  
 284 photon without sending a signal to the EOM used to  
 285 compress the photon. The model, without considering the  
 286 attenuation of the atomic cloud, is given by  $(AL * L')(\omega)$   
 287 (Fig. 4, blue line). This is provided as a reference for a  
 288 Fourier-transform limited photon with an exponentially  
 289 decaying envelope, as emitted by a single atom—the  
 290 scenario examined in the theory section.  
 291

292 To apply the analysis in the theory section for predicting  
 293 the spectrum of the compressed photon, we first rescale the  
 294 power spectrum  $|\Psi''(\omega)|^2$ , calculated from Eq. (7), with  $A$ ,  
 295 which was extracted from the previous fit. Then, we  
 296 convolve this spectrum with  $L'(\omega; \Gamma_{\text{FP}})$  to obtain the  
 297 expected compressed photon spectrum  $(A|\Psi''|^2 * L')(\omega)$ .  
 298 Figure 4 (red line) shows the modeled power spectrum  
 299 slightly deviating from the measured result (red dots),  
 300 exhibiting a lower peak coincidence rate where an absorp-  
 301 tion dip occurs in the uncompressed photon spectrum (blue  
 302 dots). We attribute this difference to the fact that our model  
 303 does not fully account for the effects imposed by the  
 304 atomic cloud.

305 *Discussion.*—By definition, spectral compression  
 306 reduces the width of a spectral distribution, resulting in  
 307 an increased photon rate and intensity at the central

308 frequency. In our experiment, we observed a bandwidth of  
 309 20(2) MHz for the initial photon, and 8(2) MHz for the  
 310 compressed photon from spectroscopy using a 2.6 MHz  
 311 cavity. This almost matches the natural  $D$  transition line-  
 312 width of 6 MHz in  $^{87}\text{Rb}$ . The maximal photon transmission  
 313 through the spectroscopic cavity is increased by a factor of  
 314 2.39(4), indicating a successful spectral compression of  
 315 narrowband photons.

316 The compressions mechanism is, in principle, lossless  
 317 since both cavity and phase modulators can have arbitrarily  
 318 low losses. In our experiment, however, we observe an  
 319 overall transmission of 22% through the compression  
 320 optics; the dispersion optics alone (PBS, QWP, dispersion  
 321 cavity) has a transmission of 72%, and the fiber-based  
 322 EOM a transmission of 30% including the fiber coupling.

323 To compare the compression method with simple passive  
 324 filtering, we calculate the transmission  $T$  of the 20 MHz  
 325 bandwidth photons produced by our source through both  
 326 bandwidth-limiting schemes, and an analyzer cavity with a  
 327 bandwidth of 6 MHz bandwidth and resonant with the  
 328 central frequency of the input photons to model an atomic  
 329 absorption process corresponding to the  $5S_{1/2} \rightarrow 5P_{1/2}$   
 330 transition in  $^{87}\text{Rb}$ . With a lossless compression system, we  
 331 find  $T = 44\%$ , while a resonant filter cavity of the same  
 332 bandwidth of 6 MHz leads to  $T = 14\%$ , illustrating the  
 333 advantage of the compression method. By replacing the  
 334 fiber-based EOM with a free-space EOM with an optical  
 335 transmission  $> 95\%$ , the compression method would  
 336 significantly surpass the transmission of a passive filter.

337 Optimal spectral compression of a photon with band-  
 338 width  $\Gamma_p$  in the cavity-based scheme is achieved if the  
 339 dispersion cavity has a bandwidth of  $0.25\Gamma_p$ . Since the  
 340 amount of spectral compression is limited by the dispersion  
 341 mechanism, dispersion engineering of structured dielectric  
 342 media [37–39] or multiple combined optical cavities may  
 343 allow us to further increase the spectral compression. This  
 344 method is not limited to the atomic system in our experi-  
 345 ment—it can be adapted to a wide range of wavelengths  
 346 and spectral widths, and therefore even allow us to match  
 347 the spectral properties to different types of quantum  
 348 systems, e.g., in a hybrid quantum network [40].

349 We thank Adrian Nugraha Utama for useful discussions  
 350 on the theoretical modeling and Jianwei Lee for his  
 351 valuable input while writing this script. This work was  
 352 **4** supported by the Ministry of Education in Singapore.

355  
 356  
 357 \*christian.kurtsiefer@gmail.com

358 **5** [1] J.I. Cirac, P. Zoller, H.J. Kimble, and H. Mabuchi,  
 359 *Phys. Rev. Lett.* **78**, 3221 (1997).  
 360 [2] H.-J. Briegel, W. Dür, J. I. Cirac, and P. Zoller, *Phys. Rev.*  
 361 *Lett.* **81**, 5932 (1998).  
 362 [3] E. Waks and C. Monroe, *Phys. Rev. A* **80**, 062330 (2009).  
 [4] H. J. Kimble, *Nature (London)* **453**, 1023 (2008).

[5] T. Wilk, S. C. Webster, A. Kuhn, and G. Rempe, *Science* **317**, 488 (2007). 363  
 364  
 [6] K. Hammerer, A. S. Sørensen, and E. S. Polzik, *Rev. Mod.*  
 365 *Phys.* **82**, 1041 (2010). 366  
 367  
 [7] M. D. Lukin, *Rev. Mod. Phys.* **75**, 457 (2003). 367  
 [8] M. Steiner, V. Leong, M. A. Seidler, A. Cerè, and C.  
 368 Kurtsiefer, *Opt. Express* **25**, 6294 (2017). 369  
 [9] M. Keller, B. Lange, K. Hayasaka, W. Lange, and H.  
 370 Walther, *Nature (London)* **431**, 1075 (2004). 371  
 [10] M. Almendros, J. Huwer, N. Piro, F. Rohde, C. Schuck, M.  
 372 Hennrich, F. Dubin, and J. Eschner, *Phys. Rev. Lett.* **103**,  
 373 213601 (2009). 374  
 [11] A. Kuhn, M. Hennrich, and G. Rempe, *Phys. Rev. Lett.* **89**,  
 375 067901 (2002). 376  
 [12] B. Srivathsan, G. K. Gulati, B. Chng, G. Maslennikov, D. N.  
 377 Matsukevich, and C. Kurtsiefer, *Phys. Rev. Lett.* **111**,  
 378 123602 (2013). 379  
 [13] J. Park, H. Kim, and H. S. Moon, *Phys. Rev. Lett.* **122**,  
 380 143601 (2019). 381  
 [14] C. Kurtsiefer, S. Mayer, P. Zarda, and H. Weinfurter, *Phys.*  
 382 *Rev. Lett.* **85**, 290 (2000). 383  
 [15] P. Michler, A. Kiraz, C. Becher, W. Schoenfeld, P. Petroff,  
 384 L. Zhang, E. Hu, and A. Imamoglu, *Science* **290**, 2282  
 385 (2000). 386  
 [16] E. Moreau, I. Robert, J. Gérard, I. Abram, L. Manin, and V.  
 387 Thierry-Mieg, *Appl Phys. Lett.* **79**, 2865 (2001). 388  
 [17] E. Meyer-Scott, N. Montaut, J. Tiedau, L. Sansoni, H.  
 389 Herrmann, T. J. Bartley, and C. Silberhorn, *Phys. Rev. A* **95**,  
 390 061803(R) (2017). 391  
 [18] C. Schuck, F. Rohde, N. Piro, M. Almendros, J. Huwer, M.  
 392 W. Mitchell, M. Hennrich, A. Haase, F. Dubin, and J.  
 393 Eschner, *Phys. Rev. A* **81**, 011802(R) (2010). 394  
 [19] L.-M. Duan, M. D. Lukin, J. I. Cirac, and P. Zoller,  
 395 *Nature (London)* **414**, 413 (2001). 396  
 [20] B. B. Blinov, D. L. Moehring, L.-M. Duan, and C. Monroe,  
 397 *Nature (London)* **428**, 153 (2004). 398  
 [21] J. McKeever, A. Boca, A. D. Boozer, R. Miller, J. R. Buck,  
 399 A. Kuzmich, and H. J. Kimble, *Science* **303**, 1992 (2004). 400  
 [22] F. Wolfgramm, X. Xing, A. Cerè, A. Predojević, A. M.  
 401 Steinberg, and M. W. Mitchell, *Opt. Express* **16**, 18145  
 402 (2008). 403  
 [23] M. Eisaman, A. André, F. Massou, M. Fleischhauer, A.  
 404 Zibrov, and M. Lukin, *Nature (London)* **438**, 837 (2005). 405  
 [24] L. Zhu, X. Guo, C. Shu, H. Jeong, and S. Du, *Appl Phys.*  
 406 *Lett.* **110**, 161101 (2017). 407  
 [25] B. Buchler, M. Hosseini, G. Hétet, B. Sparkes, and P. K.  
 408 Lam, *Opt. Lett.* **35**, 1091 (2010). 409  
 [26] B. M. Sparkes, M. Hosseini, C. Cairns, D. Higginbottom,  
 410 G. T. Campbell, P. K. Lam, and B. C. Buchler, *Phys. Rev. X*  
 411 **2**, 021011 (2012). 412  
 [27] B. H. Kolner and M. Nazarathy, *Opt. Lett.* **14**, 630 (1989). 413  
 [28] B. H. Kolner, *IEEE J. Quantum Electron.* **30**, 1951 (1994). 414  
 [29] J. Lavoie, J. M. Donohue, L. G. Wright, A. Fedrizzi, and K.  
 415 J. Resch, *Nat. Photonics* **7**, 363 (2013). 416  
 [30] M. Karpiński, M. Jachura, L. J. Wright, and B. J. Smith, **6**  
 417 *Nat. Photonics* **11**, 53 (2017). 418  
 [31] Y. Li, T. Xiang, Y. Nie, M. Sang, and X. Chen, *Sci. Rep.* **7**,  
 419 43494 (2017). 420  
 [32] B. Srivathsan, G. K. Gulati, A. Cerè, B. Chng, and C.  
 421 Kurtsiefer, *Phys. Rev. Lett.* **113**, 163601 (2014). 422



423	[33] O. Morin, M. Körber, S. Langenfeld, and G. Rempe, <i>Phys. Rev. Lett.</i> <b>123</b> , 133602 (2019).	[38] X. Li, M. Pu, X. Ma, Y. Guo, P. Gao, and X. Luo, <i>J. Phys. D</i> <b>51</b> , 054002 (2018).	432
424			433
425	[34] B. Srivathsan, G. K. Gulati, A. Cerè, B. Chng, and C. Kurtsiefer, <i>Phys. Rev. Lett.</i> <b>113</b> , 163601 (2014).	[39] J. D. Joannopoulos, S. G. Johnson, J. N. Winn, and R. D. Meade, <i>Photonic Crystals: Molding the Flow of Light</i> , 2nd ed. (Princeton University Press, Princeton, NJ, USA, 2008).	434
426			435
427	[35] H. H. Jen, <i>Phys. Rev. A</i> <b>85</b> , 013835 (2012).		436
428	[36] A. Cerè, B. Srivathsan, G. K. Gulati, B. Chng, and C. Kurtsiefer, <i>Phys. Rev. A</i> <b>98</b> , 023835 (2018).	[40] G. Kurizki, P. Bertet, Y. Kubo, K. Mølmer, D. Petrosyan, P. Rabl, and J. Schmiedmayer, <i>Proc. Natl. Acad. Sci. U.S.A.</i> <b>112</b> , 3866 (2015).	437
429			438
430	[37] E. Istrate and E. H. Sargent, <i>Rev. Mod. Phys.</i> <b>78</b> , 455 (2006).		439
431			440
			441